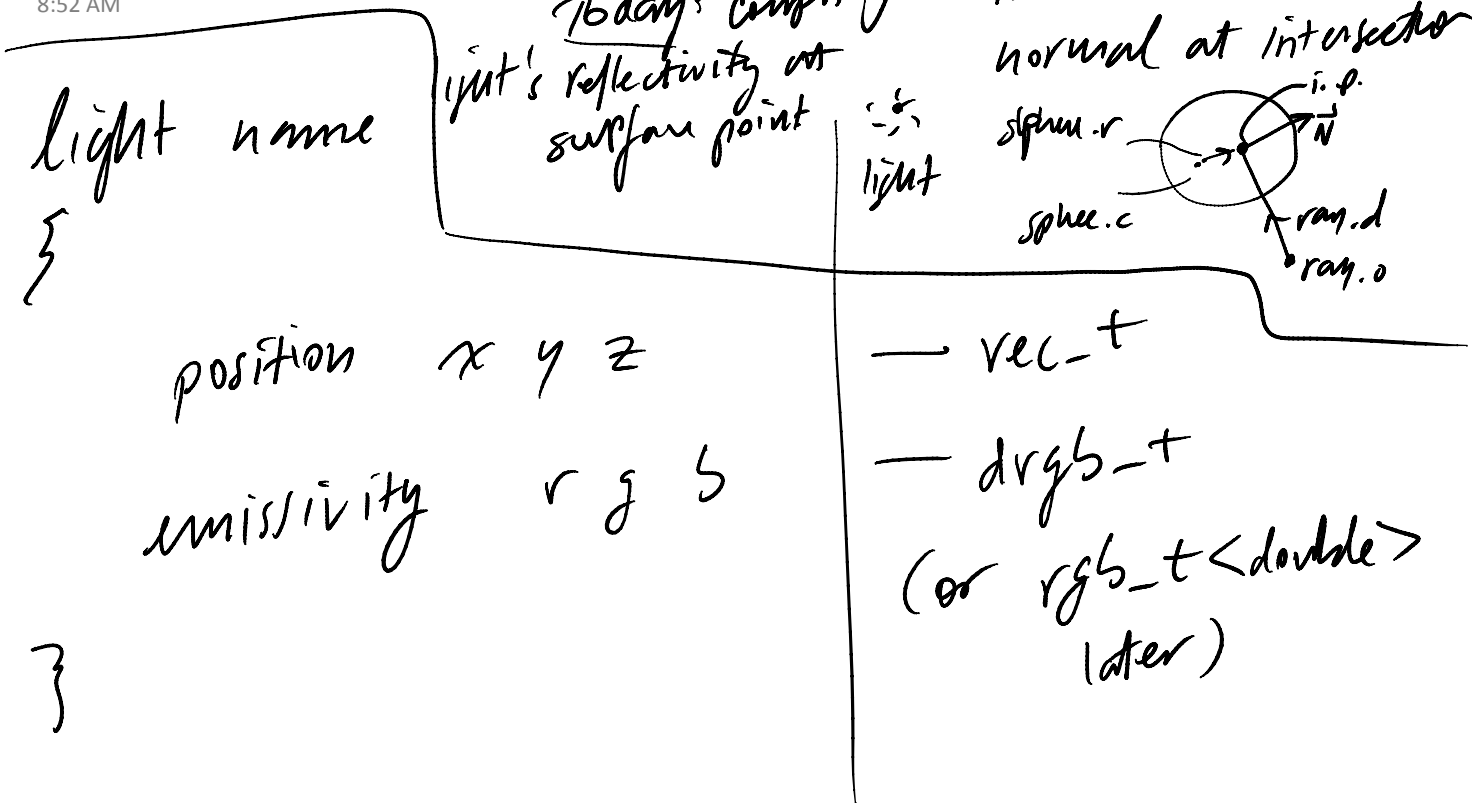


rendering equation (and light_t object)

Wednesday, October 27, 2010
8:52 AM



- emissivity is the light color, e.g. 111 is white

Rendering equation:

(at surface point, calculate intensity I)
light color (intensity, units etc)

$$I = \frac{K_a I_a}{R} + \sum_l \frac{I_l}{r} (K_d (N \cdot L) + K_s (R \cdot V)^n)$$

ambient light term,

R is the ray distance
(what you already have)

distance to light

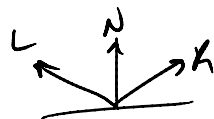
diffuse term, $K_d = obj \rightarrow getdiffuse$

specular term
 $K_s = obj \rightarrow getspecular$

where

N : surface normal
 L : direction to light
(from surface point)
(normalized)

R : L dir reflected about N



$$R = 2(N \cdot L)N - L$$

V : direction to camera
(normalized)

n : shininess term
just a double,

e.g. $n = 1.0$
 $n = 32.0$ | the larger n is, the smoother the spec. light

in code: `color = obj.getambient() / raydist;`

K_a
(surface property of object, basically the surface color)

here $I_a \approx (1, 1, 1)$

white light
(a bit of a hack)

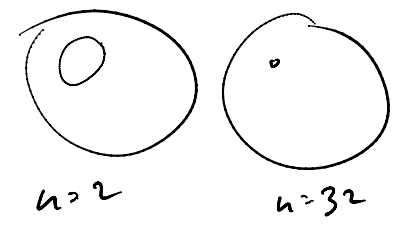
- this is called the Phong

illumination model — basically an extension of physical light transport what OpenCL uses for direct illum. (local)

1/10/07

(local)

ray tracer, by calculating
ray reflections, produces global
illumination that GL doesn't do



- Model derivation:

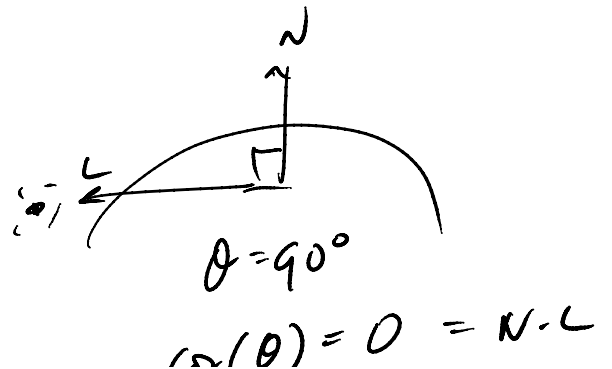
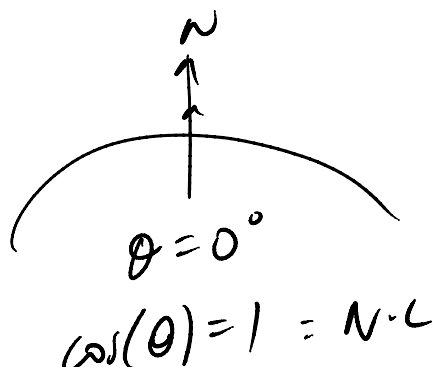
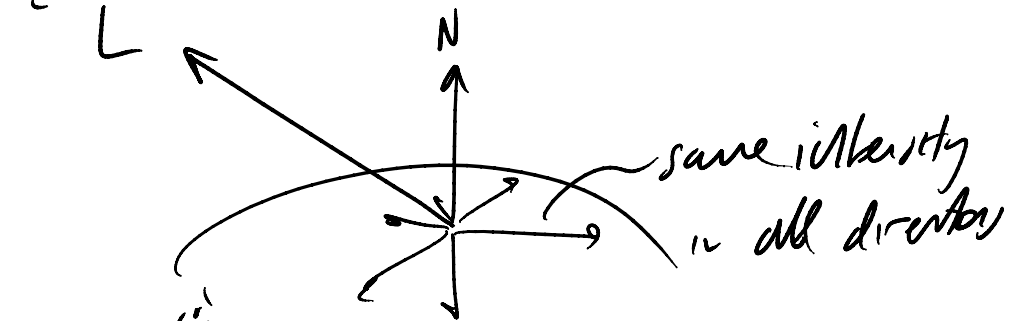
- start with diffuse scattering: light arriving at surface scattered equally in all directions (matte surface)

- Perfect diffuse reflector: reflect light according to Lambert's Law:

$$I \propto L \cdot N$$

↑
proportional to

cosine of the angle between light dir & N



(fully illuminated)

(no refraction)

$$if (0 \leq N \cdot L \leq 1)$$

surface visible by light

else

surface hidden

(only gets ambient contrib.)

- Conventional lighting model (model 1)

$$I = I_c k_d (N \cdot L)$$

↑

$o_{ij} \rightarrow$ distance

(with ambient term looks like this)

$$I = I_a k_a + \sum_l I_c k_d (N \cdot L)$$

Next refinement: scale by distance to light

Next refinement: scale by distance to "you"

in code: $L = \text{light} \rightarrow \text{get position}() - \text{last_hit}$

$r = L.\text{len}();$ // get dist to you

$L = L.\text{norm}();$ // normalize L

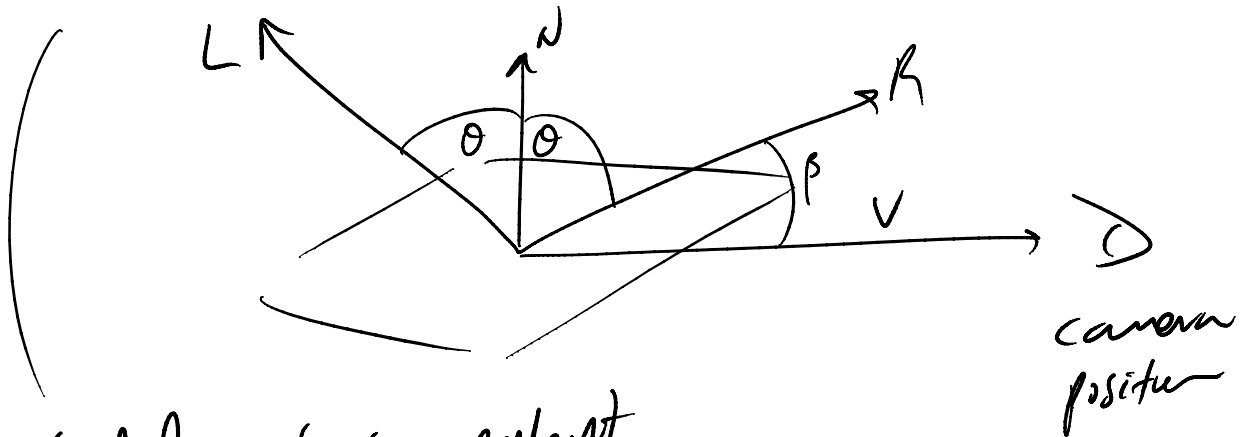
attenuated light model (model 2)

$$I = I_{\text{amb}} + \sum_l \frac{I_l (k_d (N \cdot L))}{r^2}$$

OpenCL uses
something weaker here
 $(r + k_1 + k_2)$ (?)

- still diffuse (maybe specular)

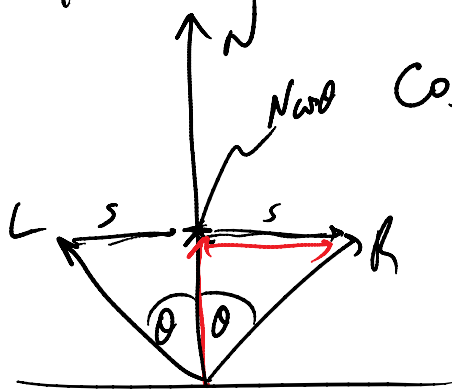
- adding in shiny component (specular)
- models the "perfect reflector" (mirror)



should only see perfect reflector if $\beta = 0$, looking down at R

- for non-perfect reflectors, intensity of reflected light falls off sharply w/ β increases

- falloff is approximated by



Normal $\cos^n \beta = (R \cdot V)^n$

$$R = N \cos \theta + s$$

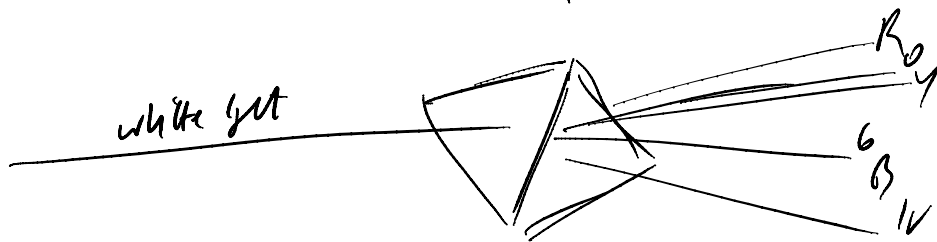
$$\text{dot } s = N \cos \theta - L \text{ via similar } \Delta$$

$$\begin{aligned}
 \text{so } f_i &= N \cos \theta + N \cos \theta - L \\
 &= 2N \cos \theta - L \\
 &= 2N(N \cdot L) - L
 \end{aligned}$$

- for real materials, amount of incident light specularly reflected depends on both angle of incidence θ and wavelength

$$I_s = I_i \omega(\lambda, \theta) \cos^4 \beta$$

↑
this requires a ray per wavelength for all wavelengths, you'd like to model



- another write:

$$I = I_a k_a + \sum_l \frac{I_l}{r} (k_d (N \cdot L) + k_r (R \cdot V)^n)$$

Blinn

$$I = I_a k_a + \sum_l \frac{I_l}{V} (k_d(N-l) + k_1(N-H)^n)$$

Blin

divisor of $L \approx V$

$$H = \frac{L+V}{\|L+V\|} \approx \frac{1}{2}(L+V)$$